Proceedings IAU Symposium No. 267, 2009

© 2009 International Astronomical Union

Brad Peterson, Rachel Somerville, & Thaisa Storchi-Bergmann eds. DOI: 00.0000/X0000000000000X

Beware of fake AGNs

G. Stasińska¹, N. Vale Asari^{1,2}, R. Cid Fernandes² for the SEAGal collaboration (Semi-Empirical Analysis of GALaxies)

¹LUTH, Observatoire de Paris, CNRS, Université Paris Diderot; Meudon, France ²Dpto de Física - CFM - Universidade Federal de Santa Catarina, Florianópolis, SC, Brazil

Abstract. In the BPT diagram, the distribution of the emission-line galaxies from the Sloan Digital Sky Survey (SDSS) evokes the wings of a seagull. Traditionally, galaxies in the right wing are considered to host AGNs. Our study of the stellar populations of SDSS galaxies showed that $\sim 1/4$ of galaxies thought to host LINERS are in fact "retired galaxies", i.e. galaxies that stopped forming stars and are ionized by hot post-AGB stars and white dwarfs (Stasińska et al. 2008). When including the galaxies that lack some of the lines needed to place them in the BPT diagram the fraction of retired galaxies is even larger (Cid Fernandes et al., 2009, arXiv:0912.1376).

Keywords. galaxies: active — galaxies: evolution — galaxies: stellar content

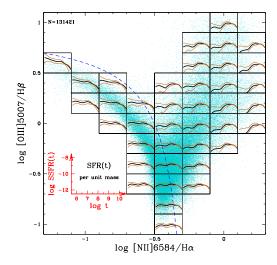


Figure 1. The BPT diagram for 131421 galaxies in the SDSS. The dashed line separates pure star forming galaxies from the rest (Stasińska et al. 2006). Superimposed are Specific Star Formation Histories obtained with the stellar population synthesis code STARLIGHT (Cid Fernandes et al. 2005). Clearly, Seyfert galaxies (upper branch of the right wing) still form stars while LINERs (lower branch) do not. Photoionization models using the radiation from the stellar populations in the LINER region are able to cover the entire BPT plane (Stasińska et al 2008). The seagull shape is the result of selection effects.

References

Cid Fernandes R., Mateus A., Sodré L., et al., 2005, MNRAS, 358, 363 Stasińska G., Cid Fernandes R., Mateus A., et al., 2006, MNRAS, 371, 972 Stasińska G., Vale Asari N., Cid Fernandes R., et al., 2008, MNRAS, 391, L29